Terry-Horton Equations

Modeling 2D Drift-Wave Turbulence with the Catherine Felce (University of Oxford), advised by Prof. Greg Hammett & Noah Mandell (Princeton Plasma Physics Laboratory)

GOALS

- Simulate 2D drift-waves with a linear driving term.
- Develop model to describe turbulence on both ITG and ETG (Ion/Electron Temperature Gradient) scales without solving for all wavenumbers in between.

BACKGROUND:

2D Drift-Wave Model

- We use the Terry-Horton equations, derived as follows ([1], [2], [3], [4]).
- Conservation of ion-guiding-center density gives:

$$\frac{\partial n_{gc}}{\partial t} + \nabla$$

$$\frac{\partial u_{gc}}{\partial t} + \boldsymbol{\nabla} \cdot [n_{gc}(\boldsymbol{V}_{\boldsymbol{E}\times\boldsymbol{B}} + u_{\parallel}\boldsymbol{\hat{z}})] = 0$$

Here u// is the parallel ion flow, and ion-guiding-center density is related to the real-space ion density in the following way due to ion-polarization effects:

$$\mathbf{n}_i = n_{gc} + n_0 \frac{\rho_s^2 e}{T_e} \boldsymbol{\nabla}_{\perp}^2 \boldsymbol{\phi}$$

The drift velocity is given by:

$$\boldsymbol{V}_{\boldsymbol{E}\times\boldsymbol{B}} = \frac{c}{B}\boldsymbol{\hat{z}}\times\boldsymbol{\nabla}\phi$$

Adding a linear drive, characterized by the delta_0 parameter, to the adiabatic (Boltzmann) electron distribution gives the following electron response:

$$\delta n_e = n_{eo} \left(1 - \delta_0 \rho_s \frac{\partial}{\partial y} \right) \frac{e\phi}{T_e}$$

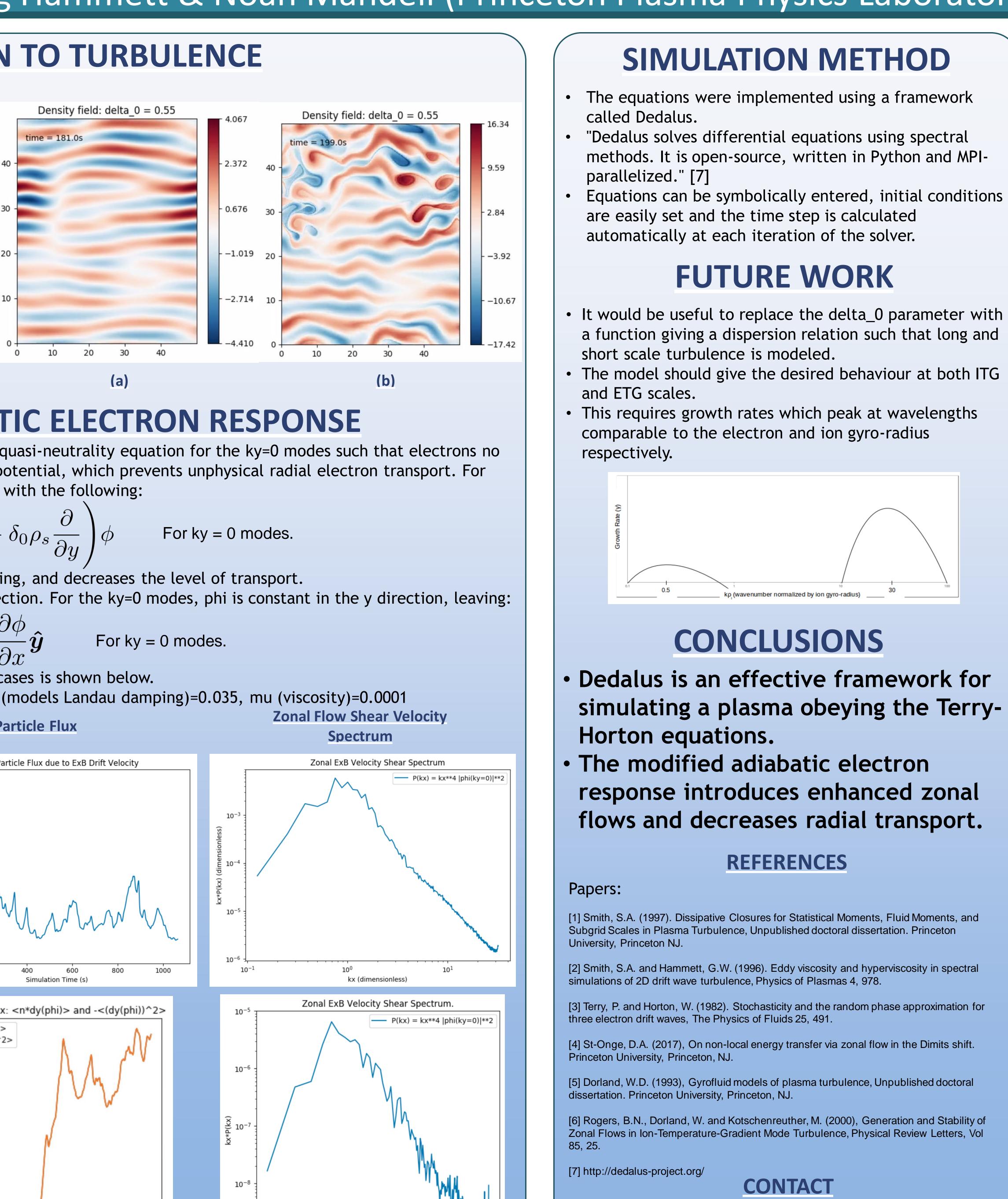
Quasi-neutrality gives the following relation between ion density and potential:

$$\tilde{n}_{gc} = \frac{n_0 e}{T_e} \left(1 - \rho_s^2 \nabla_{\perp}^2 - \delta_0 \rho_s \frac{\partial}{\partial y} \right) \phi$$

We approximate the ion-guiding-center density as a linear function of x (the radial direction), having a gradient scale length L_n, and we consider the shortscale variations of ion density. We introduce the diamagnetic drift velocity V_d, and describe u// via a Landau damping model. Adding a dissipative term involving viscosity (mu) we arrive at the following:

$$\frac{\partial \tilde{n}_{gc}}{\partial t} + \mathbf{V}_{\mathbf{E} \times \mathbf{B}} \cdot \boldsymbol{\nabla} \tilde{n}_{gc} + n_0 V_d \frac{\partial}{\partial y} \left(\frac{e\phi}{T_e} \right)$$
$$= -\alpha \frac{c_s}{L_n} \tilde{n}_{gc} + \mu \frac{c_s \rho_s^2}{L_n} \boldsymbol{\nabla}_{\perp}^2 \tilde{n}_{gc}$$

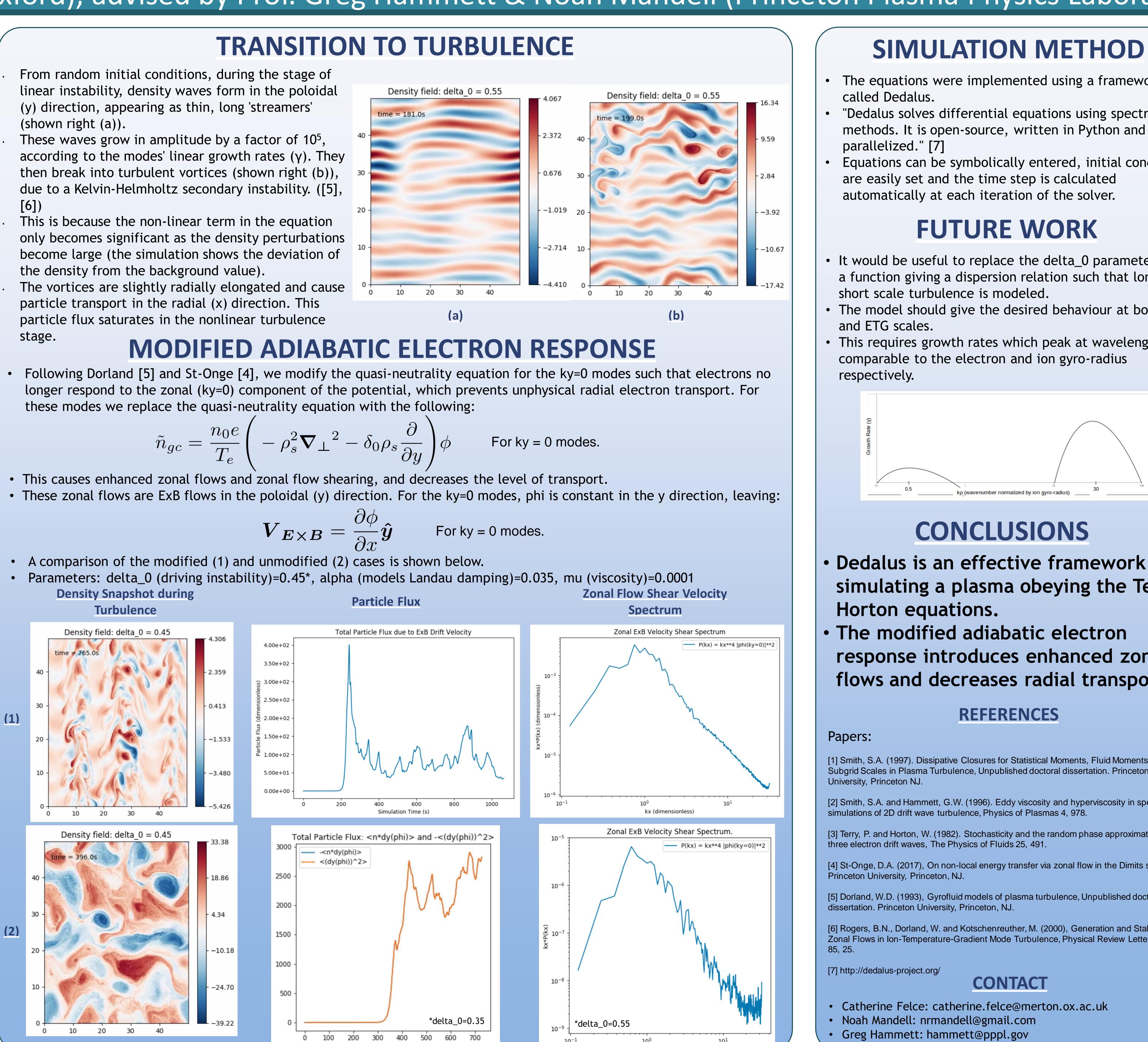
- From random initial conditions, during the stage of (y) direction, appearing as thin, long 'streamers' (shown right (a)).
- These waves grow in amplitude by a factor of 10^5 , then break into turbulent vortices (shown right (b)), [6])
- This is because the non-linear term in the equation become large (the simulation shows the deviation of
- particle transport in the radial (x) direction. This particle flux saturates in the nonlinear turbulence



$$\tilde{n}_{gc} = \frac{n_0 e}{T_e} \left(-\rho_s^2 \nabla_{\perp}^2 - \delta_0 \rho_s \frac{\partial}{\partial y} \right)$$

$${m V}_{{m E} imes {m B}} = rac{\partial \phi}{\partial x} {m \hat{y}}$$

Density Snapshot during





- methods. It is open-source, written in Python and MPI-
- Equations can be symbolically entered, initial conditions

- a function giving a dispersion relation such that long and

simulating a plasma obeying the Terry-